

Quantum Field Theory on Curved Backgrounds. I

The Euclidean Functional Integral

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Abstract We give a mathematical construction of Euclidean quantum field theory on certain curved backgrounds. We focus on generalizing Osterwalder Schrader quantization, as these methods have proved useful to establish estimates for interacting fields on flat space-times. In this picture, a static Killing vector generates translations in Euclidean time, and physical positivity is played by positivity under reflection of Euclidean time. We discuss the quantization of flows which correspond to classical space-time symmetries, and give a general set of conditions which imply that broad classes of operators in the classical picture give rise to well-defined operators on the quantum-field Hilbert space. In particular, Killing fields on spatial sections give rise to unitary groups on the quantum-field Hilbert space, and corresponding densely-defined self-adjoint generators. We construct the Schrödinger representation using a method which involves localizing certain integrals over the full manifold to integrals over a codimension-one submanifold. This method is called sharp-time localization, and implies reflection positivity.

Introduction

The present article presents a construction of a Euclidean quantum field theory on time-independent, curved backgrounds. Earlier work on field theories on curved space-time (Kay [33], Dimock [14], Bros et al. [7]) uses real-time/Lorentzian signature and algebraic techniques reminiscent of $\mathcal{P}(\varphi)_2$ theory from the Hamiltonian point of view [22]. In contrast, the present treatment uses the Euclidean functional integral [23] and Osterwalder-Schrader

quantization [38, 39]. Experience with constructive field theory on \mathbb{R}^d shows that the Euclidean functional integral provides a powerful tool, so it is interesting also to develop Euclidean functional integral methods for manifolds.

Euclidean methods are known to be useful in the study of black holes, and a standard strategy for studying black hole (BH) thermodynamics is to analytically continue time in the BH metric [10]. The present paper implies a mathematical construction of scalar fields on any static, Euclidean black hole background. The applicability of the Osterwalder-Schrader quantization procedure to curved space depends on unitarity of the time translation group and the time reflection map which we prove (theorem 2.5). The Osterwalder-Schrader construction has universal applicability; it contains the Euclidean functional integral associated with scalar boson fields, a generalization of the Berezin integral for fermions, and a further generalization for gauge fields [2]. It also appears valid for fields on Riemann surfaces [28], conformal field theory [17], and may be applicable to string theory. The present paper extends this construction to models on curved backgrounds.

Our paper has many relations with other work. Wald [42] studied metrics with Euclidean signature, although he treated the functional integral from a physical rather than a mathematical point of view. Brunetti et al [8] developed the algebraic approach (Haag-Kastler theory) for curved spacetimes and generalized the work of Dimock [13]. They describe covariant functors between the category of globally hyperbolic spacetimes with isometric embeddings, and the category of $*$ -algebras with unital injective $*$ -monomorphisms.

The examples studied in this paper—scalar quantum field theories on static space-times—have physical relevance. A first approximation to a full quantum theory (involving the gravitational field as well as scalar fields) arises from treating the sources of the gravitational field classically and independently of the dynamics of the quantized scalar fields [6]. The weakness of gravitational interactions, compared with elementary particle interactions of the standard model, leads one to believe that this approximation is reasonable. It exhibits nontrivial physical effects which are not present for the scalar field on a flat spacetime, such as the Hawking effect [25] or the Fulling-Unruh effect [41]. Density perturbations in the cosmic microwave background (CMB) are calculated using scalar field theory on certain curved backgrounds [35]. Further, Witten [45] used quantum field theory on Euclidean anti-de Sitter space in the context of the AdS/CFT correspondence [24, 36].

Some of the methods discussed here in Section 2 have been developed for the flat case in lecture courses; see [27].

Notation and conventions

We use notation, wherever possible, compatible with standard references on relativity [44] and quantum field theory [23]. We use Latin indices $a, b =$

$0 \dots d - 1$ for spacetime indices, reserving Greek indices $\mu, \nu = 1 \dots d - 1$ for spatial directions. We include in our definition of ‘Riemannian manifold’ that the underlying topological space must be paracompact (every open cover has a locally finite open refinement) and connected. The notation $L^2(M)$ is used when M is a C^∞ Riemannian manifold, and implicitly refers to the Riemannian volume measure on M , which we sometimes denote by $d\text{vol}$. Also $\mathcal{U}(\mathcal{H})$ denotes the group of unitary operators on \mathcal{H} . Let $G = I(M) = \text{Iso}(M)$ denote the isometry group, while \mathfrak{K} is its Lie algebra, the global Killing fields. For ψ a smooth map between manifolds, we use ψ^* to denote the pullback operator $(\psi^*f)(p) = f(\psi(p))$. The notation $\Delta = \Delta_M$ means the Laplace operator for the Riemannian metric on M .

1 Reflection Positivity

1.1 Analytic continuation

The Euclidean approach to quantum field theory on a curved background has advantages since elliptic operators are easier to deal with than hyperbolic operators. To obtain physically meaningful results one must presumably perform the analytic continuation back to real time. In general, Lorentzian spacetimes of interest may not be sections of 4-dimensional complex manifolds which also have Riemannian sections, and even if they are, the Riemannian section need not be unique. Thus, the general picture of extracting physics from the Euclidean approach is a difficult one where further investigation is needed.

Fortunately, for the class of spacetimes treated in the present paper (static spacetimes), the embedding within a complex 4-manifold with a Euclidean section is guaranteed, and in such a way that Einstein’s equation is preserved [11].

1.2 Time reflection

Reflection in Euclidean time plays a fundamental role in Euclidean quantum field theory, as shown by Osterwalder and Schrader [38, 39].

Definition 1.1 (Time reflection) *Let M be a Riemannian manifold. A time reflection $\theta : M \rightarrow M$ is an isometric involution which fixes pointwise a smooth codimension-one hypersurface Σ . This means that $\theta \in \text{Iso}(M)$, $\theta^2 = 1$ and $\theta(x) = x$ for all $x \in \Sigma$.*

We now discuss time reflection for static manifolds, which is the example that we will study in this paper.

Example 1.1 (Static manifolds) Suppose there exists a globally defined, static Killing field ξ . Fix a hypersurface $\Sigma \subset M$ to which ξ is orthogonal. Define a global function $t : M \rightarrow \mathbb{R}$ by setting $t = 0$ on Σ , and otherwise define $t(p)$

to be the unique number t such that $\phi_t(x) = p$ for some $x \in \Sigma$, where $\{\phi_t\}$ is the one-parameter group of isometries determined by ξ . Finally, define θ to map a point $p \in M$ to the corresponding point on the same ξ -trajectory but with $t(\theta(p)) = -t(p)$. This defines a decomposition

$$M = \Omega_- \cup \Sigma \cup \Omega_+, \quad \theta\Omega_{\pm} = \Omega_{\mp}, \quad \theta\Sigma = \Sigma. \quad (1.1)$$

In past work [28], we have considered time-reflection maps which fall outside the bounds of example 1.1 ([28] applies to compact Riemann surfaces, which cannot support Killing fields), but we will not do so here.

The time-reflection map given by a hypersurface-orthogonal Killing field is not unique, but depends on a choice of the *initial hypersurface*, which we fix. The initial hypersurface will be used to define time-zero fields. Reflection of the Euclidean time coordinate $t \rightarrow -t$ analytically continues to Hermitian conjugation of e^{-itH} .

1.3 Fundamental assumptions

Let $C = (-\Delta + m^2)^{-1}$ be the resolvent of the Laplacian, also called the *free covariance*, where $m^2 > 0$. Then C is a bounded self-adjoint operator on $L^2(M)$. For each $s \in \mathbb{R}$, the Sobolev space $H_s(M)$ is a real Hilbert space, which can be defined as completion of $C^\infty(M)$ in the norm

$$\|f\|_s^2 = \langle f, C^{-s}f \rangle. \quad (1.2)$$

We work with test functions in $H_{-1}(M)$. This is a convenient choice for several reasons: the norm (1.2) with $s = -1$ is related in a simple way to the free covariance, and further, Dimock [15] has given an elegant proof of reflection positivity for Sobolev test functions. Another motivation is as follows. Suppose we wish to prove that $\varphi(h)$ is a bounded perturbation of the free Hamiltonian H_0 for a scalar field on \mathbb{R}^d . The first-order perturbation is

$$-\langle \Omega_1, H_0 \Omega_1 \rangle = -\frac{1}{2} \int \frac{|\hat{h}(\mathbf{p})|^2}{\omega(\mathbf{p})^2} d\mathbf{p} \quad (1.3)$$

where we used $\Omega_1 = -H_0^{-1}\varphi(h)\Omega$. Existence of (1.3) is equivalent to $h \in H_{-1}(\mathbb{R}^d)$, so this is a natural condition for test functions. Therefore we choose $H_{-1}(M)$ for the generalization to curved manifolds.

The Sobolev spaces give rise to a natural rigging, or *Gelfand triple*, and various associated Gaussian measures [18, 40]. The inclusion $H_s \hookrightarrow H_{s+k}$ for $k > 0$ is Hilbert-Schmidt, so the spaces

$$H_\infty \equiv \bigcap_{s=1}^{\infty} H_s(M) \subset H_{-1}(M) \subset \bigcup_{s=-1}^{-\infty} H_s(M) \equiv H_{-\infty}$$

form a Gelfand triple, and H_∞ is a nuclear space. There is a unique Gaussian measure μ defined on the dual $H_{-\infty}$ with covariance C . This means that

$$S(f) \equiv \int_{H_{-\infty}} e^{i\Phi(f)} d\mu(\Phi) = e^{-\frac{1}{2}\langle f, Cf \rangle}, \quad f \in H_\infty.$$

Define

$$\mathcal{E} := L^2(H_{-\infty}, \mu).$$

The space \mathcal{E} is unitarily equivalent to Euclidean Fock space over $H_{-1}(M)$ (see for example [40, Theorem I.11]). The algebra generated by monomials of the form $\Phi(f_1) \dots \Phi(f_n)$ is dense in \mathcal{E} . This is a special case of a general construction discussed in the reference.

Definition 1.2 (Standard domain) *For an open set $\Omega \subseteq M$, the standard domain in \mathcal{E} corresponding to Ω is:*

$$E_\Omega = \text{span}\{e^{i\Phi(f)} : f \in H_{-1}(M), \text{supp}(f) \subset \Omega\}.$$

Let \mathcal{E}_Ω denote the closure in \mathcal{E} of E_Ω .

Definition 1.2 refers to subspaces of \mathcal{E} generated by functions supported in an open set. This includes empty products, so $1 \in \mathcal{E}_\Omega$ for any Ω . Of particular importance for Euclidean field theory is the positive-time subspace

$$\mathcal{E}_+ := \mathcal{E}_{\Omega_+},$$

where the notation Ω_+ refers to the decomposition (1.1). A linear operator on \mathcal{E} which maps $\mathcal{E}_+ \rightarrow \mathcal{E}_+$ is said to be *positive-time invariant*.

1.4 Operator induced by a diffeomorphism

We will consider the effect which diffeomorphisms of the underlying space-time manifold have on the Hilbert space operators which arise in the quantization of a classical field theory. For $f \in C^\infty(M)$ and $\psi : M \rightarrow M$ a diffeomorphism, define

$$f^\psi \equiv \psi_* f = (\psi^{-1})^* f = f \circ \psi^{-1}. \quad (1.4)$$

The reason for using ψ^{-1} here is so that Definition 1.3 gives a group representation.

Definition 1.3 (Induced operator) *Let ψ be a diffeomorphism, and $A(\Phi) = \Phi(f_1) \dots \Phi(f_n) \in \mathcal{E}$ a monomial. Define*

$$\Gamma(\psi) : A : \equiv : \Phi(f_1^\psi) \dots \Phi(f_n^\psi) : . \quad (1.5)$$

*This extends linearly to a dense domain in \mathcal{E} . We refer to $\Gamma(\psi)$ as the operator **induced by the diffeomorphism ψ** .*

Note that if ψ is an isometry, then (1.5) is equivalent to the definition $\Gamma(\psi)A \equiv \Phi(f_1^\psi) \dots \Phi(f_n^\psi)$ without Wick ordering, as follows from (1.8) below.

The induced operators $\Gamma(\psi)$ are *not* necessarily bounded on \mathcal{E} . In fact, for a general diffeomorphism ψ , the operator ψ^* may fail to be bounded on $L^2(M)$ or $H_{-1}(M)$. If the Jacobian $|d\psi|$ satisfies uniform upper and lower bounds, i.e.

$$(\exists c_1, c_2 > 0) \quad c_1 < \sup_{x \in M} |d\psi_x| < c_2. \quad (1.6)$$

then $(\psi^{-1})^*$ is bounded on $L^2(M)$, but $\Gamma(\psi)$ may still be unbounded on \mathcal{E} , because the operator norm of $\Gamma(\psi)$ on the degree- n subspace of \mathcal{E} may fail to have a limit as $n \rightarrow \infty$. In this situation, $\Gamma(\psi)$ is to be regarded as a densely-defined unbounded operator whose domain includes all finite particle vectors.

If $(\psi^{-1})^*$ is a contraction on $H_{-1}(M)$, then $\Gamma(\psi)$ is a contraction on \mathcal{E} (in particular, bounded). A special case of this is $\psi \in \text{Iso}(M)$, which implies that $\Gamma(\psi)$ is unitary and $\|\Gamma(\psi)\|_{\mathcal{E}} = 1$.

Lemma 1.1 (Naturalness property) *Let $\psi : M \rightarrow M$ be a diffeomorphism, and consider the pullback ψ^* acting on $L^2(M)$, with its Hermitian adjoint $(\psi^*)^\dagger$. Then*

$$\det(d\psi) = 1 \Leftrightarrow (\psi^*)^\dagger = (\psi^{-1})^* \Leftrightarrow \psi \text{ is volume-preserving}. \quad (1.7)$$

Furthermore,

$$\psi \in \text{Iso}(M) \Leftrightarrow \Gamma(\psi) \in \mathcal{U}(\mathcal{E}) \Leftrightarrow [\psi^*, \Delta] = 0 \Leftrightarrow [\psi^*, C] = 0. \quad (1.8)$$

The last part of (1.8) follows from [32, Theorem III.6.5], while the rest of the statements in (1.7) and (1.8) are standard calculations. It follows that Γ restricts to a unitary representation of $G = \text{Iso}(M)$ on \mathcal{E} .

For an open set $\Omega \subset M$, define

$$\text{Iso}(M, \Omega) = \{\psi \in \text{Iso}(M) : \psi(\Omega) \subset \Omega\},$$

and similarly $\text{Diff}(M, \Omega)$. These are not subgroups of $\text{Diff}(M)$ but they are semigroups under composition. If $\psi \in \text{Diff}(M, \Omega)$ we say ψ preserves Ω .

Lemma 1.2 (Presheaf property) *Let $\psi : U \rightarrow V$ be a diffeomorphism, where U, V are open sets in M . Let E_U, E_V be the corresponding standard domains (cf. Definition 1.2). Then*

$$\Gamma(\psi)E_U = E_V.$$

In particular, if $\psi : M \rightarrow M$ preserves $\Omega \subset M$, then $\Gamma(\psi)$ preserves the corresponding subspace $\mathcal{E}_\Omega \subset \mathcal{E}$.

For maps $\psi : U \rightarrow V$ which are subset inclusions $U \subseteq V$, lemma 1.2 asserts that the association $U \rightarrow \mathcal{E}_U$ is a *presheaf*. It also follows from lemma 1.2 that the mappings $U \rightarrow \mathcal{E}_U$ and $\psi \rightarrow \Gamma(\psi)$ define a covariant functor from the category of open subsets of M with invertible, smooth maps between them into the category of Hilbert spaces and densely defined operators.

Lemma 1.2 implies that if $\psi(\Omega_+) \subset \Omega_+$ then $\Gamma(\psi)$ is positive-time invariant. This is necessary but not sufficient for $\Gamma(\psi)$ to have a quantization. A sufficient condition is that $\Gamma(\psi)$ and $\Theta\Gamma(\psi)^\dagger\Theta$ both preserve \mathcal{E}_+ , where $\Theta = \Gamma(\theta)$, as shown by Theorem 2.1.

1.5 Continuity results

Lemma 1.3 (Sobolev continuity) *For the free covariance $C = (-\Delta + m^2)^{-1}$,*

$$\{f_1, \dots, f_n\} \longmapsto A(\Phi) := \Phi(f_1) \dots \Phi(f_n) \in \mathcal{E}$$

is continuous $(H_{-1})^n \rightarrow \mathcal{E}$, where we take the product of the Sobolev topologies on $(H_{-1})^n$.

Proof Since Φ is linear, it is sufficient to show that $\|A(\Phi)\|_{\mathcal{E}}$ is bounded by $\text{const.} \prod_i \|f_i\|_{-1}$. As a consequence of the Gaussian property of the measure $d\mu_C$, one needs only bound the linear case. But

$$\|\Phi(f)\|_{\mathcal{E}} = \left| \int (\Phi(f)\Phi(f)) d\mu_C \right|^{1/2} = \|f\|_{-1}. \quad (1.9)$$

Theorem 1.1 (Strong continuity) *Let $\{\psi_n\}$ be a sequence of orientation-preserving isometries which converge to ψ in the compact-open topology. Then $\Gamma(\psi_n) \rightarrow \Gamma(\psi)$ in the strong operator topology on $\mathcal{B}(\mathcal{E})$.*

The proof of theorem 1.1 follows standard arguments in analysis. Let us give a sense of how it is to be used. If all the elements of a certain one-parameter group of isometries ψ_t are such that $\Gamma(\psi_t)$ have bounded quantizations, then $t \rightarrow \hat{\Gamma}(\psi_t)$ defines a one-parameter group of operators on \mathcal{H} (the quantum-field Hilbert space). In this situation, Theorem 1.1 justifies the application of Stone's theorem. This picture is to be developed in section 2.

1.6 Reflection positivity

Definition 1.4 *With θ as in Definition 1.1, let $\Theta = \Gamma(\theta)$ be the induced reflection on \mathcal{E} . A measure μ on $H_{-\infty}$ is said to be reflection positive if*

$$\int \overline{\Theta(F)} F d\mu \geq 0 \quad \text{for all } F \in \mathcal{E}_+. \quad (1.10)$$

A bounded operator T on $L^2(M)$ is said to be reflection positive if

$$\text{supp } f \subseteq \Omega_+ \Rightarrow \langle f, \Theta T f \rangle_{L^2(M)} \geq 0. \quad (1.11)$$

Reflection positivity for the measure μ is equivalent to the following inequality for operators on $\mathcal{E} = L^2(d\mu)$:

$$0 \leq \Pi_+ \Theta \Pi_+$$

where $\Pi_+ : \mathcal{E} \rightarrow \mathcal{E}_+$ is the canonical projection.

A Gaussian measure with mean zero and covariance C is reflection positive iff C is reflection positive in the operator sense, eqn. (1.11). An equivalent condition is that for any finite sequence $\{f_i\}$ of real functions supported in Ω_+ , the matrix $M_{ij} = \exp \langle f_i, \theta C f_j \rangle$ has no negative eigenvalues.

For Riemannian manifolds which possess an isometric involution whose fixed-point set has codimension one, there is a simple potential-theoretic proof of reflection positivity [12]. The relation between reflection positivity and operator monotonicity under change of boundary conditions for the Laplacian was discovered in [21]. A different proof of reflection positivity on curved spaces was given by Dimock [15], based on Nelson's proof using the Markov property [37]. We give a third proof later in this paper based on our sharp-time localization theorem. The result is summarized as follows.

Theorem 1.2 (Reflection positivity) *Let M be a Riemannian manifold with a time reflection as in Definition 1.1. Then the covariance $C = (-\Delta + m^2)^{-1}$ and its associated Gaussian measure are reflection positive.*

2 Osterwalder-Schrader Quantization and the Feynman-Kac Formula

The Osterwalder-Schrader construction is a standard feature of quantum field theory. It begins with a "classical" Euclidean Hilbert space \mathcal{E} and leads to the construction of a Hilbert space $\mathcal{H} = \Pi \mathcal{E}_+$, which is the projection Π of the Euclidean space \mathcal{E}_+ . It also yields a quantization map $T \mapsto \hat{T}$ from a classical operator T on \mathcal{E} to a quantized operator \hat{T} acting on \mathcal{H} . In this section we review this construction, dwelling on the quantization of bounded operators T on \mathcal{E} that may yield a bounded or an unbounded quantization \hat{T} , as well as the quantization of an unbounded operator T on \mathcal{E} . We give a variation of the previously unpublished treatment in [27], adapted to curved space-time.

2.1 The Hilbert space

Define a bilinear form (A, B) on \mathcal{E}_+ by

$$(A, B) = \langle \Theta A, B \rangle_{\mathcal{E}} \quad \text{for } A, B \in \mathcal{E}_+. \quad (2.1)$$

Using self-adjointness of Θ on \mathcal{E} , one can show that this form is sesquilinear,

$$(B, A) = \int \overline{\Theta B} A \, d\mu = \left(\int B \overline{\Theta A} \, d\mu \right)^* = \overline{(A, B)}. \quad (2.2)$$

If θ is not an isometry, then Θ is non-unitary in which case Osterwalder-Schrader quantization is not possible. Therefore, it is essential that $\theta \in \text{Iso}(M)$. The form (2.1) is degenerate, and has an infinite-dimensional kernel which we denote \mathcal{N} . Therefore (2.1) determines a nondegenerate inner product $\langle, \rangle_{\mathcal{H}}$ on $\mathcal{E}_+/\mathcal{N}$, making the latter a pre-Hilbert space.

Definition 2.1 (Hilbert space) *The (Osterwalder-Schrader) physical Hilbert space \mathcal{H} is the completion of $\mathcal{E}_+/\mathcal{N}$, with inner product $\langle, \rangle_{\mathcal{H}}$. Let $\Pi : \mathcal{E}_+ \rightarrow \mathcal{H}$ denote the natural quotient map, a contraction mapping from $A \in \mathcal{E}_+$ to $\hat{A} := \Pi A$. There is an exact sequence:*

$$0 \longrightarrow \mathcal{N} \xrightarrow{\text{incl.}} \mathcal{E}_+ \xrightarrow{\Pi} \mathcal{H} \longrightarrow 0 .$$

2.2 Quantization of operators

Assume that T is a densely defined, closable operator on \mathcal{E} with domain $\mathcal{D} \subset \mathcal{E}$. Define $T^+ := \Theta T^* \Theta$, and assume there exists a subdomain $\mathcal{D}_0 \subset \mathcal{D} \cap \mathcal{E}_+$ on which T^+ is defined and for which both

$$T: \mathcal{D}_0 \rightarrow \mathcal{E}_+, \text{ and } T^+: \mathcal{D}_0 \rightarrow \mathcal{E}_+ . \quad (2.3)$$

Theorem 2.1 (Condition for quantization) *Assume that $\hat{\mathcal{D}}_0 := \Pi(\mathcal{D}_0)$ is dense in \mathcal{H} . Condition (2.3) ensures that T has a quantization \hat{T} with domain $\hat{\mathcal{D}}_0$. Furthermore \hat{T}^* is defined, \hat{T} has a closure, and on $\hat{\mathcal{D}}_0$, we have:*

$$\hat{T}^* = \widehat{T^+} . \quad (2.4)$$

Proof First, we check that \hat{T} is well-defined. Suppose $A \in \mathcal{N} \cap \mathcal{D}_0$. Let $B \in \mathcal{E}_+$ range over a set of vectors in the domain of $\Theta T^* \Theta$ such that the image of this set under Π is dense in \mathcal{H} . Then

$$0 = \langle (\Theta T^* \Theta B)^\wedge, \hat{A} \rangle_{\mathcal{H}} = \langle T^* \Theta B, A \rangle_{\mathcal{E}} = \langle \Theta B, TA \rangle_{\mathcal{E}} = \langle \hat{B}, \widehat{TA} \rangle_{\mathcal{H}} .$$

Thus $TA \in \mathcal{N}$, and hence T is well-defined on $\mathcal{D}_0/\mathcal{D}_0 \cap \mathcal{N}$. To check (2.4) is a routine calculation.

The main content of Theorem 2.1 can be expressed as a commutative diagram. For bounded transformations, Theorem 2.1 simply means that if $T : \mathcal{E}_+ \rightarrow \mathcal{E}_+$ and the dotted arrow in the following diagram is well-defined, then so are the two solid arrows:

$$\begin{array}{ccccccc} 0 & \longrightarrow & \mathcal{N} & \xrightarrow{\text{incl.}} & \mathcal{E}_+ & \xrightarrow{\Pi} & \mathcal{H} \longrightarrow 0 \\ & & \downarrow T & & \downarrow \Theta T^* \Theta & & \downarrow \hat{T} \\ 0 & \longrightarrow & \mathcal{N} & \xrightarrow{\text{incl.}} & \mathcal{E}_+ & \xrightarrow{\Pi} & \mathcal{H} \longrightarrow 0 \end{array}$$

Lemma 2.1 (Contraction property) *Let T be a bounded transformation on \mathcal{E} such that T and $\Theta T^* \Theta$ each preserve \mathcal{E}_+ . Then \hat{T} is a bounded transformation on \mathcal{H} and*

$$\|\hat{T}\|_{\mathcal{H}} \leq \|T\|_{\mathcal{E}}. \quad (2.5)$$

Proof This proceeds by the multiple reflection method [23].

We now discuss some examples of operators satisfying the hypotheses of Theorem 2.1.

Theorem 2.2 (Self-adjointness) *Let U be unitary on \mathcal{E} , and $U(\mathcal{E}_+) \subset \mathcal{E}_+$. If $U^{-1}\Theta = \Theta U$ then U admits a quantization \hat{U} and \hat{U} is self-adjoint. (Do not assume U^{-1} preserves \mathcal{E}_+).*

Proof The operator $\Theta U^* \Theta = \Theta^2 U = U$ preserves \mathcal{E}_+ , so Theorem 2.1 $\Rightarrow U$ has a quantization \hat{U} . Self-adjointness of \hat{U} follows from eqn. (2.4).

Theorem 2.3 (Unitarity) *Let U be unitary on \mathcal{E} , and $U^{\pm 1}(\mathcal{E}_+) \subset \mathcal{E}_+$. If $[U, \Theta] = 0$ then U admits a quantization \hat{U} and \hat{U} is unitary.*

Proof The operator $\Theta U^* \Theta = U^* = U^{-1}$ preserves \mathcal{E}_+ by assumption, so U has a quantization. Also, $\Theta(U^{-1})^* \Theta = U$ preserves \mathcal{E}_+ , so U^{-1} also has a quantization. Obviously, the quantization of U^{-1} is the inverse of \hat{U} . Eqn. (2.4) implies that the adjoint of \hat{U} is the quantization of $\Theta U^* \Theta = U^* = U^{-1}$.

Examples of operators satisfying the conditions of Theorems 2.2 and 2.3 come naturally from isometries on M with special properties. We now discuss two classes of isometries, which give rise to self-adjoint and unitary operators as above.

Example 2.1 (Reflected Isometries) An element $\psi \in \text{Iso}(M)$ is said to be a **reflected** isometry if

$$\psi^{-1} \circ \theta = \theta \circ \psi. \quad (2.6)$$

If additionally $\psi(\Omega_+) \subseteq \Omega_+$ then Theorem 2.2 implies that $\hat{\Gamma}(\psi) : \mathcal{H} \rightarrow \mathcal{H}$ exists and is self-adjoint. If ψ satisfies (2.6) then so does ψ^{-1} ; hence if $\psi^{-1}(\Omega_+) \subseteq \Omega_+$, then $\Gamma(\psi^{-1})$ has a quantization and $\hat{\Gamma}(\psi^{-1})$ is the inverse of $\hat{\Gamma}(\psi)$.

Example 2.2 (Reflection-Invariant Isometries) A **reflection-invariant** isometry is an element $\psi \in \text{Iso}(M)$ that commutes with time-reflection, $\psi\theta = \theta\psi$. It follows that $[\Gamma(\psi), \Theta] = 0$. If ψ and ψ^{-1} both preserve Ω_+ then $\Gamma(\psi^{\pm 1})\mathcal{E}_+ \subset \mathcal{E}_+$, and Theorem 2.3 implies that $\hat{\Gamma}(\psi) : \mathcal{H} \rightarrow \mathcal{H}$ is unitary. The set of reflection-invariant isometries form a subgroup of the full isometry group.

2.3 Quantization domains

Quantization domains are subsets of Ω_+ which give rise to dense domains in \mathcal{H} after quantization. This is important for the analysis of unbounded operators on \mathcal{H} . For example, an isometry which satisfies (2.6) may only map a proper subset $\mathcal{O} \subset \Omega_+$ into Ω_+ , and in this case $\Gamma(\psi)$ is only defined on a non-dense subdomain of \mathcal{E}_+ . If \mathcal{O} is a quantization domain, then $\Pi\mathcal{E}_{\mathcal{O}}$ may still be dense in \mathcal{H} , and can serve as a domain of definition for $\hat{\Gamma}(\psi)$.

Definition 2.2 *A quantization domain is a subspace $\Omega \subset \Omega_+$ with the property that $\Pi(\mathcal{E}_{\Omega})$ is dense in \mathcal{H} .*

Example 2.3 Perhaps the simplest quantization domain is a half-space lying at times greater than $T > 0$,

$$\mathcal{O}_{+,T} = \{x \in \mathbb{R}^d : x_0 > T\}. \quad (2.7)$$

Let $\mathcal{D}_{+,T} = E_{\mathcal{O}_{+,T}} = \Gamma(\psi_T)E_+$ where $\psi_T(x, t) = (x, t + T)$; then $\Pi(\mathcal{D}_{+,T})$ is dense in \mathcal{H} , as follows from Theorem 2.4.

Theorem 2.4 generalizes (2.7) to curved spacetimes, and also allows one to replace the simple half-space $\mathcal{O}_{+,T}$ with a more general connected subset of Ω_+ .

Theorem 2.4 (Construction of quantization domains) *Suppose $\psi \in \text{Iso}(M, \Omega_+)$, i.e. $\mathcal{O} := \psi(\Omega_+) \subset \Omega_+$. If $[\Gamma(\psi), \Theta] = 0$ or $\Gamma(\psi)\Theta = \Theta\Gamma(\psi^{-1})$ (i.e. ψ is reflection-invariant or reflected) then \mathcal{O} is a quantization domain.*

Proof By lemma 1.2, we have

$$E_{\mathcal{O}} = \Gamma(\psi)E_+. \quad (2.8)$$

Let $\hat{C} \in \mathcal{H}$ be orthogonal to every vector $\hat{A} \in \Pi(\mathcal{E}_{\mathcal{O}})$. Choose $B \in \mathcal{E}_+$ and let $A := \Gamma(\psi)B \in \mathcal{E}_{\mathcal{O}}$. Then

$$0 = \langle \hat{C}, \hat{A} \rangle_{\mathcal{H}} = \langle \hat{C}, \Pi(\Gamma(\psi)B) \rangle_{\mathcal{H}} = \langle \Theta C, \Gamma(\psi)B \rangle_{\mathcal{E}}.$$

Since $\Gamma(\psi)^{-1}$ is unitary on \mathcal{E} , apply it to the inner product to yield

$$\langle \Gamma(\psi^{-1})\Theta C, B \rangle_{\mathcal{E}} = 0 \quad (\forall B \in \mathcal{E}_+).$$

Therefore $\Gamma(\psi^{-1})\Theta C$ is orthogonal (in \mathcal{E}) to the entire subspace \mathcal{E}_+ .

First, suppose that $[\Gamma(\psi^{-1}), \Theta] = 0$. Then we infer

$$0 = \langle \Theta\Gamma(\psi^{-1})C, B \rangle_{\mathcal{E}} = \langle \hat{\Gamma}(\psi^{-1})\hat{C}, \hat{B} \rangle_{\mathcal{H}} \quad (\forall \hat{B} \in \Pi(E_+)),$$

i.e. $\hat{C} \in \ker \hat{\Gamma}(\psi^{-1})$. Therefore,

$$(\Pi(\mathcal{E}_{\mathcal{O}}))^{\perp} = \ker \hat{\Gamma}(\psi^{-1}). \quad (2.9)$$

Since $[\Gamma(\psi^{-1}), \Theta] = 0$ then Theorem 2.3 implies that $\hat{\Gamma}(\psi)$ is unitary, hence the kernel of $\hat{\Gamma}(\psi^{-1})$ is trivial and $\Pi(\mathcal{E}_{\mathcal{O}})$ is dense in \mathcal{H} . We have thus completed the proof in this case.

Now, assume that $\Gamma(\psi)\Theta = \Theta\Gamma(\psi^{-1})$. Example 2.1 implies that $\hat{\Gamma}(\psi)$ exists and is self-adjoint on \mathcal{H} , and moreover (by the same argument used above),

$$(\Pi(\mathcal{E}_{\mathcal{O}}))^{\perp} = \ker \hat{\Gamma}(\psi).$$

If $\psi = \psi_t$ where $\{\psi_s\}$ is a one-parameter group of isometries, and if $\hat{\Gamma}(\psi_t)$ is a strongly continuous semigroup then by Stone's theorem, $\hat{\Gamma}(\psi_t) = e^{-tK}$ for K self-adjoint. Since e^{-tK} clearly has zero kernel, the proof is also complete in the second case.

Corollary 2.1 *The set $\mathcal{O}_{+,T}$ is a quantization domain.*

The problem of characterizing all quantization domains appears to be open.

2.4 Construction of the Hamiltonian and ground state

Theorem 2.5 (Time-translation semigroup) *Let $\xi = \partial/\partial t$ be the time-translation Killing field on the static spacetime M . Let the associated one-parameter group of isometries be denoted $\phi_t : M \rightarrow M$. For $t \geq 0$, $U(t) = \Gamma(\phi_t)$ has a quantization, which we denote $R(t)$. Further, $R(t)$ is a well-defined one-parameter family of self-adjoint operators on \mathcal{H} satisfying the semigroup law.*

Proof Lemma 1.1 implies that $U(t)$ is unitary on \mathcal{E} , and it is clearly a one-parameter group. Also,

$$\phi_t \circ \theta = \theta \circ \phi_{-t}$$

and $U(t)\mathcal{E}_+ \subset \mathcal{E}_+$ for $t \geq 0$, so this is a *reflected isometry*; see Example 2.1. Theorem 2.2 implies $R(t) = \hat{U}(t)$ is a self-adjoint transformation on \mathcal{H} for $t \geq 0$, which satisfies the group law

$$R(t)R(s) = R(t+s) \quad \text{for } t, s \geq 0$$

wherever it is defined.

Theorem 2.6 (Hamiltonian and ground state) *$R(t)$ is a strongly continuous contraction semigroup, which leaves invariant the vector $\Omega_0 = \hat{1}$. There exists a densely defined, positive, self-adjoint operator H such that*

$$R(t) = \exp(-tH), \quad \text{and } H\Omega_0 = 0.$$

Thus Ω_0 is a quantum-mechanical ground state.

Proof It is immediate that $R(t)\Omega_0 = \Omega_0$. The contraction property $R(t) \leq I$ follows from the multiple reflection method, as explained in [23]. The remaining statements are consequences of Stone's theorem.

The operator H is the quantum mechanical generator (in the Euclidean picture) of translations in the direction ξ . When $\xi = \partial/\partial t$, then H is called the *Hamiltonian*.

2.5 Feynman-Kac theorem

Theorem 2.7 (Feynman-Kac) *Let $\hat{A}, \hat{B} \in \mathcal{H}$, and let H be the Hamiltonian constructed in Theorem 2.6. Each matrix element of the heat kernel e^{-tH} is given by a Euclidean functional integral,*

$$\langle \hat{A}, e^{-tH} \hat{B} \rangle_{\mathcal{H}} = \int \overline{\Theta A} U(t) B \, d\mu(\Phi). \quad (2.10)$$

The right-hand side of (2.10) is the Euclidean path integral [16] of quantum field theory. Mark Kac' method [30, 31] for calculating the distribution of the integral $\int_0^T v(X_t) dt$, where v is a function defined on the state space of a Markov process X , gives a rigorous version of Feynman's work, valid at imaginary time.

In the present setup, (2.10) requires no proof, since the functional integral on the right-hand side is how we defined the matrix element on the left-hand side. However, some work is required (even for flat spacetime, $M = \mathbb{R}^d$) to see that the Hilbert space and Hamiltonian given by this procedure take the usual form arising in physics. This is true, and was carried out for \mathbb{R}^d by Osterwalder and Schrader [38] and summarized in [23, Ch. 6].

Since H is positive and self-adjoint, the heat kernels can be analytically continued $t \rightarrow it$. We therefore define the *Schrödinger group* acting on \mathcal{H} to be the unitary group

$$R(it) = e^{-itH}.$$

Given a time-zero field operator, action of the Schrödinger group then defines the corresponding real-time field.

For flat spacetimes in $d \leq 3$ it is known [23] that Theorem 2.7 has a generalization to non-Gaussian integrals, i.e. interacting quantum field theories:

$$\begin{aligned} \langle \hat{A}, e^{-tH_V} \hat{B} \rangle_{\mathcal{H}} &= \left\langle \Theta A, \exp \left(- \int_0^t dt' \int d\mathbf{x} V(\Phi(\mathbf{x}, t')) \right) B_t \right\rangle_{\mathcal{E}} \\ &= \int \overline{\Theta A} e^{-S_{0,t}^V} B_t \, d\mu(\Phi). \end{aligned} \quad (2.11)$$

Construction the non-Gaussian measure (2.11) in finite volume can presumably be completed by a straightforward extension of present methods, while the infinite-volume limit seems to require a cluster expansion. Work is in progress to address these issues for curved spacetimes.

2.6 Quantization of subgroups of the isometry group

Physics dictates that after quantization, a spacetime symmetry with p parameters should correspond to a unitary representation of a p -dimensional Lie group acting on \mathcal{H} . The group of spacetime symmetries for Euclidean quantum field theory should be related to the group for the real-time theory by analytic continuation; this was shown for flat spacetime by Klein

and Landau [34]. For curved spacetimes, no such construction is known, and due to the intrinsic interest of such a construction, we give further details, and show that the methods already discussed in this paper suffice to give a unitary representation of the purely spatial symmetries on \mathcal{H} .

Example 2.2 introduced reflection-invariant isometries. We now discuss an important subclass of these, the purely spatial isometries, which are guaranteed to have well-defined quantizations. We continue to assume we have a static manifold M with notation as in Example 1.1. There is a natural subgroup G_{space} of $G = \text{Iso}(M)$ consisting of isometries which map each spatial section into itself. We term these **purely spatial** isometries. The classic constructions [29] of finite-volume interactions in two dimensions work on a cylinder $M = S^1 \times \mathbb{R}$, in which case G_{space} is the subgroup of $\text{Iso}(S^1 \times \mathbb{R})$ corresponding to rotations around the central axis.

Define the centralizer of θ in G as usual:

$$C_G(\theta) \equiv \{\phi \in G \mid \phi\theta = \theta\phi\}.$$

Note that

$$G_{\text{space}} \subset \text{Iso}(M, \Omega_+) \cap C_G(\theta), \quad (2.12)$$

and although the right hand side of (2.12) is not a subgroup, this is a compact way of expressing that the elements are positive-time invariant and null-invariant. Since $G_{\text{space}} \subset G$ as a Lie subgroup, $\mathfrak{g}_{\text{sp}} = \text{Lie}(G_{\text{space}})$ is a subalgebra of \mathfrak{K} , the Lie algebra of global Killing fields.

Consider the restriction of the unitary representation Γ to the subgroup G_{space} . By a standard construction, the derivative $D\Gamma$ is a unitary Lie algebra representation of \mathfrak{g}_{sp} on \mathcal{E} , for which \mathcal{E}_+ is an invariant subspace. The latter property is crucial; if \mathcal{E}_+ is not an invariant subspace for an operator, then that operator does not have a quantization.

As with many aspects of Osterwalder-Schrader quantization, a commutative diagram is helpful:

$$\begin{array}{ccc} G_{\text{space}} & \xrightarrow{\Gamma} & \mathcal{U}(\mathcal{E}) \\ \text{Lie} \downarrow & & \downarrow \text{Lie} \\ \mathfrak{g}_{\text{sp}} & \xrightarrow{D\Gamma} & \mathfrak{u}(\mathcal{E}) \end{array} \quad (2.13)$$

Note that $\mathcal{U}(\mathcal{E})$ is an infinite-dimensional Lie group. Further, there are delicate analytic questions involving the domains of the symmetric operators in $\mathfrak{u}(\mathcal{E})$. In the present paper we investigate only the algebraic structure.

By Theorem 2.3, each one-parameter unitary group $U(t)$ on \mathcal{E}_+ coming from a one-parameter subgroup of G_{space} has a well-defined quantization $\hat{U}(t)$ which is a unitary group on \mathcal{H} . The methods of Section 1.5 establish strong continuity for these unitary groups, so their generators are densely-defined self-adjoint operators as guaranteed by Stone's theorem.

Suppose that $[X, Y] = Z$ for three elements $X, Y, Z \in \mathfrak{g}_{\text{sp}}$. Let $\hat{X} : \mathcal{H} \rightarrow \mathcal{H}$ be the quantization of $D\Gamma(X)$, and similarly for Y and Z . Our

assumptions guarantee that $[D\Gamma(X), D\Gamma(Y)] = D\Gamma(Z)$ is null-invariant, therefore we have

$$[\hat{X}, \hat{Y}] = \hat{Z}, \quad (2.14)$$

valid on the domain of vectors in \mathcal{H} where the expressions are defined.

One-parameter subgroups coming from G_{space} always admit unitary representations on \mathcal{H} , but for other subgroups of G , the analogous theory is much more subtle. Since any element of \mathfrak{K} is a vector field acting on functions as a differential operator, it is local (does not change supports) and hence positive-time invariant, so quantization applied directly to infinitesimal generators may be possible. There, one runs into delicate domain issues. A discussion of the domains of some self-adjoint operators obtained by this procedure was given in Section 2.3, and some variant of this could possibly be used to treat the domains of the quantized generators.

When applied to isometry groups, Osterwalder-Schrader quantization of operators involves the procedure of taking the derivative of a representation, applied to the infinite-dimensional group $\mathcal{U}(\mathcal{E})$. Thus, it is not surprising that it is functorial, adding to its intrinsic mathematical interest. These connections are likely to lead to an interesting new direction in representation theory, especially for noncompact groups.

3 Variation of the Metric

3.1 Metric dependence of matrix elements in quantum field theory

We wish to obtain rigorous analytic control over how quantum field theory on a curved background depends upon the metric.

Definition 3.1 (Stable family) *Let M_λ denote the Riemannian manifold diffeomorphic to $\mathbb{R} \times S$, endowed with the product metric*

$$ds_\lambda^2 := dt^2 + G_{\mu\nu}(\lambda) dx^\mu dx^\nu, \quad (3.1)$$

where $G(\lambda)$ is a metric on S , and $G_{\mu\nu}(\lambda)$ depends smoothly on $\lambda \in \mathbb{R}$. We refer to a family $\{M_\lambda\}_{\lambda \in \mathbb{R}}$ satisfying these properties as a stable family. We denote the full metric (3.1) as $g(\lambda)$ or g_λ .

For a stable family, it is clearly possible to choose Ω_\pm, Σ in a way that is independent of λ . Let t denote the coordinate which is defined so that $t|_\Sigma = 0$ and $\xi = \partial/\partial t$. Then the data $(\Omega_-, \Sigma, \Omega_+, \xi, t)$ is constant in λ .

However, the Hilbert spaces $L^2(M_\lambda)$, the covariance

$$C(\lambda) = C_\lambda := (-\Delta_{g(\lambda)} + m^2)^{-1},$$

and the test function space $H_{-1}(M_\lambda)$ all depend upon λ , as does the Gaussian measure described in section 1.3. These dependences create many subtleties in the quantization procedure. In particular, the usual theory of smooth or analytic families of bounded operators does not apply to the

family of operators $\lambda \rightarrow C(\lambda)$, because if $\lambda \neq \lambda'$ then $C(\lambda)$ and $C(\lambda')$ act on different Hilbert spaces. It is clearly of interest to have some framework in which we can make sense out of the statement “ $\lambda \rightarrow C(\lambda)$ is smooth.” More generally, we would like a framework to analyze the λ -dependence of the Osterwalder-Schrader quantization.

Our approach to this set of problems is based on the observation that, for a stable family, there exist test functions $f : M \rightarrow \mathbb{R}$ which are elements of $H_{-1}(M_\lambda)$ for all λ . For example,

$$C_c^\infty(M) \subset \mathcal{H}_{-1} := \bigcap_{\lambda \in \mathbb{R}} H_{-1}(M_\lambda). \quad (3.2)$$

Such test functions can be used to give meaning to formally ill-defined expressions such as $\partial C_\lambda / \partial \lambda$. To give meaning to the naive expression

$$\frac{\partial C_\lambda}{\partial \lambda} f := \lim_{\epsilon \rightarrow 0} \frac{1}{\epsilon} (C_\lambda f - C_{\lambda+\epsilon} f), \quad (3.3)$$

we must specify the topology in which the limit is to be taken. Suppose that $f \in C_c^\infty$ as before. A natural choice is the topology of $L^2(M_\lambda)$, but some justification is necessary in the noncompact case. Clearly $C_\lambda f \in L^2(M_\lambda)$, but it is not clear that $C_{\lambda+\epsilon} f$ also determines an element of $L^2(M_\lambda)$. After all, the covariance operators are nonlocal, and $C_{\lambda+\epsilon} f$ generally does not have compact support (unless of course M itself is compact).

In order that the limit (3.3) can be taken in the topology of $L^2(M_\lambda)$, it is necessary and sufficient that $\exists \epsilon_1 > 0$ such that $C_{\lambda+\epsilon} f \in L^2(M_\lambda)$ for all $\epsilon < \epsilon_1$. In other words, the limit (3.3) makes sense iff $F(\epsilon) \equiv \int_M |C_{\lambda+\epsilon} f|^2 \sqrt{|g_\lambda|} dx < \infty$ for all $\epsilon < \epsilon_1$. Since obviously $F(0) < \infty$, it suffices to show $F(\epsilon)$ is continuous at $\epsilon = 0$. If we write the expressions in terms of coordinate charts and assume $f > 0$, then we can translate the problem into one of classical analysis. Indeed,

$$F(\epsilon) = \int_M dx \sqrt{|g_\lambda(x)|} \left(\int_{\text{supp } f} dy \sqrt{|g_{\lambda+\epsilon}(y)|} C_{\lambda+\epsilon}(x, y) f(y) \right)^2. \quad (3.4)$$

Thus the condition for differentiability of $F(\epsilon)$ at $\epsilon = 0$ becomes one of “differentiating under the integral,” which can be treated by standard methods. The overall conclusion: if $F(\epsilon)$ is continuous at $\epsilon = 0$, then (3.3) makes sense. Anticipating what is to come, this condition implies that (3.7) also makes sense.

We now return to the study of the full quantum theory on M_λ . Define

$$\mathcal{E}_\lambda := L^2(d\mu_\lambda)$$

where $d\mu_\lambda$ is the unique Gaussian probability measure associated to $C(\lambda)$ by Minlos' theorem.¹ If $f \in \mathcal{H}_{-1}$, then

$$A_{f,\lambda} = : e^{-i\Phi(f)} :_{C(\lambda)} \quad (3.5)$$

defines a canonical element of \mathcal{E}_λ for each λ . Then

$$\langle A_{f,\lambda}, A_{g,\lambda} \rangle_{\mathcal{E},\lambda} = \exp \left(\langle f, C(\lambda)g \rangle_{L^2(M_\lambda)} \right). \quad (3.6)$$

Lemma 3.1 (Smoothness of covariance) *Assume that $\{M_\lambda\}_{\lambda \in \mathbb{R}}$ is a stable family. Then $\langle f, C(\lambda)g \rangle_{L^2(M_\lambda)}$ is a smooth function of λ , for any $f, g \in C_c^\infty(M)$.*

Proof The integral $\langle f, C(\lambda)g \rangle_{L^2(M_\lambda)} = \int_M \bar{f} C_\lambda g \sqrt{|g_\lambda|} dx$ is localized over the support of f , which is compact. The dominated convergence theorem shows that we can interchange $\partial/\partial\lambda$ with the integral.

It follows immediately that the matrix element (3.6) on \mathcal{E} of the canonical elements $A_{f,\lambda}$ and $A_{g,\lambda}$ is a smooth function of the parameter λ .

When we change λ , the measure $d\mu_\lambda$ follows a path in the space of all Gaussian measures. This change in the measure can be controlled through operator estimates on the covariance. Using formula 9.1.33 from [23, p. 208] we have:

$$\frac{d}{d\lambda} \int A d\mu_\lambda = \frac{1}{2} \int (\Delta_{dC/d\lambda} A) d\mu_\lambda. \quad (3.7)$$

In particular, if $C(\lambda)$ is smooth then so is $\int A d\phi_{C(\lambda)}$. Here we must interpret $dC/d\lambda$ as in the discussion following (3.3).

The null space \mathcal{N}_λ of OS quantization also depends on the metric, as we discuss presently. When it is necessary to distinguish the time direction, we denote local coordinates by $x = (\mathbf{x}, t)$. The subspace of \mathcal{N}_λ corresponding to monomials in the field is canonically isomorphic to the space of test functions f such that²

$$\int_M f(\mathbf{x}, -t) (C_\lambda f)(\mathbf{x}, t) \sqrt{|g_\lambda(\mathbf{x})|} dx = 0. \quad (3.8)$$

All of the quantities in the integrand (3.8) which depend on λ do so smoothly. Assuming the applicability of dominated convergence arguments similar to those used above, it should be possible to show that \mathcal{N}_λ varies continuously in the Hilbert Grassmannian, but we do not address this here.

For each λ , the Osterwalder-Schrader theory gives unambiguously a quantization

$$\mathcal{H}(\lambda) \equiv \overline{\mathcal{E}_{+,\lambda}/\mathcal{N}_\lambda}.$$

¹ As before, $E_{+,\lambda} = \text{span}\{e^{i\Phi(f)} \mid f \in H_{-1}(M_\lambda), \text{supp}(f) \subset \Omega_+\}$, with completion

$$\mathcal{E}_{+,\lambda} = \overline{E_{+,\lambda}}.$$

Also define E_λ to be the (incomplete) linear span of $e^{i\Phi(f)}$ for $f \in H_{-1}(M_\lambda)$.

² For integrals such as this one, we can factorize the Laplacian as in Sec. 4.

Theorem 3.1 (Smoothness of matrix elements in \mathcal{H}) *Assume that $\{M_\lambda\}_{\lambda \in \mathbb{R}}$ is a stable family. Define the canonical element $A_{f,\lambda}$ as in (3.5). Then*

$$\lambda \rightarrow \langle \hat{A}_{f,\lambda}, R_\lambda(t) \hat{A}_{g,\lambda} \rangle_{\mathcal{H}(\lambda)}$$

is smooth.

Proof Calculate $\langle \hat{A}, R_\lambda(t) \hat{B} \rangle_{\mathcal{H}(\lambda)} = \exp \langle \theta f, (C_\lambda h) \circ \phi_{\lambda,t}^{-1} \rangle$, where $\phi_{\lambda,t}$ is the time t map of the Killing field $\partial/\partial t$ on the spacetime M_λ . Since f has compact support, the dominated convergence theorem applies to the integral $\langle \theta f, (C_\lambda h) \circ \phi_{\lambda,t}^{-1} \rangle$.

One class of examples which merits further consideration is the class formulated on $M = \mathbb{R}^{d+1}$ with $ds^2 = dt^2 + g(\lambda)_{ij} dx^i dx^j$, $i, j = 1 \dots d$. Assume that $G(\lambda)_{ij}$ depends analytically on $\lambda \in \mathbb{C}$, and to order zero it is the flat metric δ_{ij} . Theorem 3.1 implies that the matrix elements of H have a well-defined series expansion about $\lambda = 0$, and we know that precisely at $\lambda = 0$ they take their usual flat-space values.

3.2 Stably symmetric variations

It is of interest to extend the considerations of the previous section to the quantizations of symmetry generators. For this we continue to consider variations of an ultrastatic metric, as in equation (3.1). One important aspect of the quantization that is generally *not* λ -invariant is the symmetry structure of the Riemannian manifold. We assume $M = \mathbb{R} \times M'$, where M' is a Riemannian manifold with metric $g_{\mu\nu}(\lambda)$. In this section we study a special case in which the perturbation does not break the symmetry. Let \mathfrak{K}_λ denote the algebra of global Killing fields on $(M', g(\lambda))$. In certain very special cases we may have the following.

Definition 3.2 (Stable symmetry) *The family of metrics $\lambda \rightarrow g(\lambda)$ is said to be **stably symmetric** over the subinterval $I \subset \mathbb{R}$ if for each $\lambda \in I$, there exists a basis $\{\xi_i(\lambda) : 1 \leq i \leq n\}$ of \mathfrak{K}_λ , and the family of bases can be chosen in such a way that $\lambda \rightarrow \xi_i(\lambda)$ is smooth $\forall i$.*

Equivalently, the condition of stable symmetry is that $\mathfrak{K}_\lambda = KF(M_\lambda)$ gives a rank n vector bundle over \mathbb{R} (or some subinterval thereof) and we have chosen a complete set $\{\xi_i : i = 1 \dots n\}$ of smooth sections.

Example 3.1 (curvature variation) The most general constant-curvature hyperbolic metric on \mathbb{H} has arc length

$$ds = \frac{c}{\Im(z)} |dz| \tag{3.9}$$

and curvature $-c^{-2}$. Consider the spacetime $\mathbb{R} \times \mathbb{H}(c)$ where $\mathbb{H}(c)$ is the upper half-plane with metric (3.9). Variation of the curvature parameter c satisfies the assumptions of definition 3.2.

Example 3.2 (ADM mass, charge, etc.) Many spacetimes considered in physics seem to have the property of stable symmetry under variation of parameters, at least for certain ranges of those parameters. For the Euclidean continuation of the Reissner-Nordström black hole, where λ plays the role of either mass m or charge e , one may observe that the assumptions of definition 3.2 hold. However, the Euclidean RN metrics are not ultrastatic as was assumed above. Therefore, it would be interesting to extend the analysis of this section to static metrics of the form $F(\lambda, x)dt^2 + G(\lambda, x)dx^2$, where x is a $d - 1$ dimensional coordinate.

For each i, λ , the Killing field $\xi_i(\lambda)$ gives rise to a one-parameter group of isometries on M , which we denote by $\phi_{i, \lambda, x} \in \text{Iso}(M)$, where $x \in \mathbb{R}$ is the flow parameter. These flows act on the spatial section of M for each fixed time; they are *purely spatial isometries* in the sense considered above. Therefore, the map

$$T_i(\lambda, x) = \Gamma(\phi_{i, \lambda, x}) : \mathcal{E} \longrightarrow \mathcal{E}. \quad (3.10)$$

is positive-time invariant, null-invariant, and has a unitary quantization

$$\hat{T}_i(\lambda, x) : \mathcal{H} \longrightarrow \mathcal{H}. \quad (3.11)$$

None of the following constructions depend on i , so for the moment we fix i and suppress it in the notation. Since each $T(\lambda, x)$ depends on a Killing field ξ , the first step is to determine how the Killing fields vary as a function of the metric. Since the Killing fields are solutions to a first-order partial differential equation, one possible method of attack could proceed by exploiting known regularity properties of solutions to that equation. If one were to pursue that, some simplification may be possible due to the fact that a Killing field is completely determined by its first-order data at a point. We obtain a more direct proof.

The T operators depend on the Killing field through its associated one-parameter flow. For each fixed λ , the construction gives a one-parameter subgroup (in particular, a curve) in G_{space} . If we vary $\lambda \in [a, b]$, we have a free homotopy between two paths in G_{space} . Each cross-section of this homotopy, such as $\lambda \rightarrow \phi_{\lambda, x}(p)$ with the pair (x, p) held fixed, describes a continuous path in a particular spatial section of M .

Theorem 3.2 *Assume stable symmetry and define $T(\lambda, x)$ as in (3.10). Then for each x (held fixed), the map*

$$\lambda \longmapsto \hat{T}(\lambda, x) \in \mathcal{U}(\mathcal{H})$$

is a strongly continuous operator-valued function of λ .

Proof First, we claim that $\lambda \rightarrow \phi_{\lambda, x}$ is continuous in the compact-open topology. The latter follows from standard regularity theorems for solutions of ODEs, since we have assumed $\lambda \rightarrow \xi(\lambda)$ is smooth, and $\phi_{\lambda, x}(p)$ is the solution curve of the differential operator $\xi(\lambda)_p$. Theorem 1.1 implies that

$\Gamma(\phi_{\lambda,x}) \in \mathcal{U}(\mathcal{E})$ is strongly continuous with respect to λ . By theorem 2.1, the embedding of bounded operators on \mathcal{E} into $\mathcal{B}(\mathcal{H})$ is norm-continuous. Composing these continuous maps gives the desired result.

4 Sharp-time Localization

The goal of this section is to establish an analog of [23, Theorem 6.2.6] for quantization in curved space, and to show that the Hilbert space of Euclidean quantum field theory may be expressed in terms of data local to the zero-time slice. This is known as *sharp-time localization*. We first define the type of spacetime to which our results apply.

Definition 4.1 *A quantizable static spacetime is a complete, connected Riemannian manifold M with a globally defined (smooth) Killing field ξ which is orthogonal to a codimension one hypersurface $\Sigma \subset M$, such that the orbits of ξ are complete and each orbit intersects Σ exactly once.*

Under the assumptions for a quantizable static spacetime, but with Lorentz signature, Ishibashi and Wald [26] have shown that the Klein-Gordon equation gives sensible classical dynamics, for sufficiently nice initial data. These assumptions guarantee that we are in the situation of Definition 1.1.

The main difficulty in establishing sharp-time localization comes when trying to prove the analog of formula (6.2.16) of [23] in the curved space case, which would imply that the restriction to \mathcal{E}_0 of the quantization map is surjective. The proof given in [23] relies on the formula (6.2.15) from Prop. 6.2.5, and it is the latter formula that we must generalize.

4.1 Localization on flat spacetime

The Euclidean propagator on \mathbb{R}^d is given explicitly by the momentum representation

$$C(x; y) = C(x - y) = \frac{1}{(2\pi)^d} \int_{\mathbb{R}^d} \frac{1}{p^2 + m^2} e^{-ip \cdot (x-y)} dp,$$

for $x, p \in \mathbb{R}^d$. Let $f = f(\mathbf{x})$ denote a function on \mathbb{R}^{d-1} , and define

$$f_t(\mathbf{x}, t') = f(\mathbf{x}) \delta(t - t').$$

Theorem 4.1 (Flat-space localization) *Let $M = \mathbb{R}^d$ with the standard Euclidean metric. Then*

$$\langle f_t, Cg_s \rangle_{L^2(\mathbb{R}^d)} = \left\langle f, \frac{1}{2\mu} e^{(t-s)\mu} g \right\rangle_{L^2(\mathbb{R}^{d-1})}$$

where μ is the operator with momentum-space kernel $\mu(\mathbf{p}) = (\mathbf{p}^2 + m^2)^{1/2}$.

4.2 Splitting the Laplacian on static spacetimes

Consider a quantizable static space-time M , defined in Definition 4.1. Use Latin indices a, b , etc. to run from 0 to $d - 1$ and Greek indices $\mu, \nu = 1 \dots d - 1$. Denote the spatial coordinates by

$$\mathbf{x} = (x^1, \dots, x^{d-1}) = (x^\mu),$$

and set $t = x^0$. Write g in manifestly static form,

$$g_{ab} = \begin{pmatrix} F & 0 \\ 0 & G_{\mu\nu} \end{pmatrix}, \quad \text{with inverse } g^{ab} = \begin{pmatrix} 1/F & 0 \\ 0 & G^{\mu\nu} \end{pmatrix}. \quad (4.1)$$

where F and G depend only on \mathbf{x} , and not on $t = x^0$. It is then clear that

$$\mathcal{G} := \det(g_{ab}) = FG, \quad \text{where } G = \det(G_{\mu\nu}). \quad (4.2)$$

It follows that $g^{0\nu} = g^{\mu 0} = 0$, and $g^{00} = F^{-1} = g_{00}^{-1}$, does not depend upon time. Using the formula, $\Delta f = \mathcal{G}^{-1/2} \partial_a (\mathcal{G}^{1/2} g^{ab} \partial_b f)$, the Laplacian on M may be seen to be

$$\Delta_M = \frac{1}{F} \partial_t^2 + Q, \quad \text{where} \quad (4.3)$$

$$Q := \frac{1}{\sqrt{\mathcal{G}}} \partial_\mu (\sqrt{\mathcal{G}} G^{\mu\nu} \partial_\nu). \quad (4.4)$$

The operator Q is related to the Laplacian Δ_Σ for the induced metric on Σ . Applying the product rule to (4.3) yields

$$Q = \frac{1}{2} \partial_\alpha (\ln F) G^{\alpha\beta} \partial_\beta + \Delta_\Sigma. \quad (4.5)$$

Note that a formula generalizing (4.5) to “warped products” appears in Bertola et.al. [5].

In order that the operator $\mu = (-Q + m^2)^{1/2}$ exists for all $m^2 > 0$, we require that $-Q$ is a positive, self-adjoint operator on an appropriately-defined Hilbert space. The correct Hilbert space is

$$\mathcal{K}_\Sigma := L^2(\Sigma, \sqrt{\mathcal{G}} dx). \quad (4.6)$$

Here $\sqrt{\mathcal{G}} dx$ denotes the Borel measure on Σ which has the indicated form in each local coordinate system, and $\mathcal{G} = FG$ as in eq. (4.2).

Spectral theory of the operator $-Q$ considered on \mathcal{K}_Σ is mathematically equivalent to that of the “wave operator” A defined by Wald [42, 43] and Wald and Ishibashi [26]. In those references, the Klein-Gordon equation has the form $(\partial_t^2 + A)\phi = 0$. The relation between Wald’s notation and ours is that $Q = -(1/F)A - m^2$, and Wald’s function V is our $F^{1/2}$. As pointed out by Wald, we have the following,

Theorem 4.2 (Q is symmetric and negative) *Let (M, g_{ab}) be a quantizable static spacetime. Then $-Q$ is a symmetric, positive operator on the domain $C_c^\infty(\Sigma) \subset \mathcal{K}_\Sigma$.*

Proof It is easy to see that Q is symmetric on $C_c^\infty(\Sigma)$ with the metric of \mathcal{K}_Σ ; it remains to show $-Q \geq 0$ on the same domain. Using (4.4), the associated quadratic form is

$$\begin{aligned} \langle f, (-Q)f \rangle_{\mathcal{K}_\Sigma} &= - \int \bar{f} \frac{1}{\sqrt{\mathcal{G}}} \partial_\mu \left(\sqrt{\mathcal{G}} G^{\mu\nu} \partial_\nu f \right) \sqrt{\mathcal{G}} dx \\ &= \int \|\nabla f\|_G^2 \sqrt{\mathcal{G}} dx \geq 0. \end{aligned}$$

where we used integration by parts to go from the first line to the second.

4.3 Hyperbolic space

It is instructive to calculate Q in the explicit example of \mathbb{H}^d , often called *Euclidean AdS* in the physics literature because its analytic continuation is the Anti-de Sitter spacetime. The metric is

$$ds^2 = r^{-2} \sum_{i=0}^{d-1} dx_i^2, \quad r = x_{d-1}.$$

The hyperbolic Laplacian in d dimensions is (see for instance [4]):

$$\Delta_{\mathbb{H}^d} = (2-d)r \frac{\partial}{\partial r} + r^2 \Delta_{\mathbb{R}^d}. \quad (4.7)$$

Any vector field $\partial/\partial x_i$ where $i \neq d-1$ is a static Killing field. We have set up the coordinates so that it is convenient to define $t = x_0$ as before, and we can quantize in the t direction.

Comparing (4.4) with (4.7), we find that $F = r^{-2}$ and

$$Q = (2-d)r \frac{\partial}{\partial r} + r^2 \sum_{i=1}^{d-1} \frac{\partial^2}{\partial x_i^2} = -r \frac{\partial}{\partial r} + \Delta_{\mathbb{H}^{d-1}}, \quad (4.8)$$

which matches (4.5) perfectly. We return to this example spacetime in Appendix A, where we calculate its Green function, and discuss the analytic continuation.

4.4 Curved space localization

To generalize Theorem 4.1 to curved space, choose static coordinates \mathbf{x}, t near the time-zero slice Σ . If $f = f(\mathbf{x})$ is a function on the slice Σ , we define

$$f_t(\mathbf{x}, t') = f(\mathbf{x}) \delta(t - t'),$$

which is a distribution on the patch of M covered by this coordinate chart. For the moment, we assume that this coordinate patch is the region of interest. By equation (4.4), we infer that the integral kernel $\mathcal{C}(x, y)$ of the

operator $C = (-\Delta + m^2)^{-1}$ is time-translation invariant, so that we may write

$$\mathcal{C}(x, y) = \mathcal{C}(\mathbf{x}, \mathbf{y}, x_0 - y_0).$$

In order to apply spectral theory to Q , we choose a self-adjoint extension of the symmetric operator constructed by theorem 4.2. For definiteness, we may choose the Friedrichs extension, but any ambiguity inherent in the choice of a self-adjoint extension will not enter into the following analysis. We denote the self-adjoint extension also by Q , which is an unbounded operator on \mathcal{K}_Σ . The following is a generalization of Theorem 4.1 to curved space.

Theorem 4.3 (Localization of sharp-time integrals) *Let M be a quantizable static spacetime (definition 4.1). Then:*

$$\langle f_t, Cg_s \rangle_M = \left\langle f, \left(F^{1/2} \frac{e^{-|t-s|\omega}}{2\omega} F^{1/2} \right) g \right\rangle_{\mathcal{K}_\Sigma}, \quad (4.9)$$

where $\mu = (-Q + m^2)^{1/2}$ and $\omega = \left(\sqrt{F} \mu^2 \sqrt{F} \right)^{1/2}$. Hence C is reflection positive on $L^2(M)$.

Proof Because M was assumed to be a quantizable static spacetime, $F = \langle \xi, \xi \rangle_\Sigma \geq 0$. Moreover, if $F(p) = 0$ then $\xi_p = 0$, for any $p \in \Sigma$. A non-trivial Killing field cannot vanish on an open set, so the zero-set of F has measure zero in Σ . From this we infer that multiplication by the function F^{-1} defines a (possibly-unbounded) but densely-defined self-adjoint multiplication operator on \mathcal{K}_Σ .

For simplicity of notation, assume f is real-valued. Perform a partial Fourier transform with respect to the time variable:

$$\langle f_t, Cg_s \rangle_M = \int f(\mathbf{x}) \left(\frac{1}{2\pi} \int dE \frac{e^{iE(t-s)}}{F^{-1}E^2 - Q + m^2} g \right) (\mathbf{x}) \sqrt{\mathcal{G}} dx. \quad (4.10)$$

Define $\mu := (-Q + m^2)^{1/2}$, where the square root is defined through the spectral calculus on \mathcal{K}_Σ . As a consequence of theorem 4.2, μ and ω are positive, self-adjoint operators on \mathcal{K}_Σ . The integrand of (4.10) contains the operator:

$$\frac{e^{iE(t-s)}}{F^{-1}E^2 + \mu^2} = \frac{e^{iE(t-s)}}{F^{-1/2} (E^2 + F^{1/2} \mu^2 F^{1/2}) F^{-1/2}} = F^{1/2} \frac{e^{iE(t-s)}}{E^2 + \omega^2} F^{1/2}.$$

We next establish that ω is invertible. Since $\mu^2 > \epsilon I$, where $\epsilon > 0$, we have

$$\omega^2 = \sqrt{F} \mu^2 \sqrt{F} > \epsilon F$$

and therefore,

$$\omega^{-2} < \left(\sqrt{F} \mu^2 \sqrt{F} \right)^{-1} < \frac{1}{\epsilon F}.$$

Since $1/F$ is a densely defined operator on \mathcal{K}_Σ , it follows that ω^2 (hence ω) is invertible. For $\lambda > 0$,

$$\int \frac{e^{iE\tau}}{E^2 + \lambda^2} dE = \frac{\pi e^{-|\tau|\lambda}}{\lambda}. \quad (4.11)$$

Decompose the operator ω according to its spectral resolution, with $\omega = \int \lambda dP_\lambda$ and $I = \int dP_\lambda$ the corresponding resolution of the identity, and apply (4.11) in this decomposition to conclude

$$\int \frac{e^{iE(t-s)}}{F^{-1}E^2 + \mu^2} dE = F^{1/2} \frac{\pi e^{-|t-s|\omega}}{\omega} F^{1/2} \quad (4.12)$$

Inserting (4.12) into (4.10) gives

$$\begin{aligned} \langle f_t, Cg_s \rangle_M &= \int_\Sigma \left(F^{1/2} f \right) (\mathbf{x}) \left(\frac{e^{-|t-s|\omega}}{2\omega} (F^{1/2} g) \right) (\mathbf{x}) \sqrt{\mathcal{G}} d\mathbf{x} \\ &= \left\langle f, F^{1/2} \frac{e^{-|t-s|\omega}}{2\omega} F^{1/2} g \right\rangle_{\mathcal{K}_\Sigma}, \end{aligned} \quad (4.13)$$

also demonstrating reflection positivity.

The operator ω^2 may be calculated explicitly if the metric is known, and is generally not much more complicated than Q . For example, using the conventions of sec. 4.3, one may calculate ω^2 for \mathbb{H}^d :

$$\omega^2 = - \sum_{i=1}^{d-1} \partial_i^2 + d r^{-1} \partial_r + (m^2 - d) r^{-2}.$$

For \mathbb{H}^2 , the eigenvalue problem $\omega^2 f = \lambda f$ becomes a second-order ODE which is equivalent to Bessel's equation. The two linearly independent solutions are

$$r^{3/2} J_{\frac{1}{2}\sqrt{4m^2+1}}(r\sqrt{\lambda}) \quad \text{and} \quad r^{3/2} Y_{\frac{1}{2}\sqrt{4m^2+1}}(r\sqrt{\lambda}).$$

The spectrum of ω^2 on \mathbb{H}^2 is then $[0, +\infty)$.

Given a function f on Σ , we obtain a distribution f_t supported at time t as follows:

$$f_t(x, t') = f(x) \delta(t - t').$$

It may appear that this is not well-defined because it depends on a coordinate. However, given a static Killing vector, the global time coordinate is fixed up to an overall shift by a constant, which we have determined by the choice of an orthogonal hypersurface where $t = 0$. Thus a pair (p, t) where $p \in \Sigma$ and $t \in \mathbb{R}$ uniquely specify a point in M .

Theorem 4.4 (Localization of \mathcal{H}) *Let M be a quantizable static spacetime. Then the vectors $\exp(i\Phi(f_0))$ lie in \mathcal{E}_+ , and quantization maps the span of these vectors isometrically onto \mathcal{H} .*

Proof Since \mathcal{E}_+ is the *closure* of the set E_+ of vectors $\exp(i\Phi(f))$ with $\text{supp}(f) \subset \Omega_+$, it follows that any sequence in \mathcal{E}_+ which converges in the topology of \mathcal{E} has its limit in \mathcal{E}_+ . The L^2 norm in \mathcal{E} ,

$$\int \left| e^{i\Phi(f)} - e^{i\Phi(g)} \right|^2 d\mu_C(\Phi) = 2(1 - e^{-\frac{1}{2}\|f-g\|^{-1}}),$$

is controlled in terms of the norm $\|\cdot\|_{-1}$ on Sobolev space, which is the space of test functions. This will give us the first part of the theorem.

If $t > 0$, then there exists a sequence of smooth test functions $\{g_n\}$ with compact, positive-time support such that

$$\lim_{n \rightarrow \infty} g_n = f_t$$

in the Sobolev topology, hence $\exp(i\Phi(f_t)) \in \mathcal{E}_+$. Define the *time- t subspace* $\mathcal{E}_t \subset \mathcal{E}_+$ to be the subspace generated by vectors of the form $\exp(i\Phi(f_t))$. By taking the $t \rightarrow 0$ limit, we see that $\exp(i\Phi(f_0)) \in \mathcal{E}_+$ and the first part is proved.

It is straightforward to see that the quantization map $\Pi(A) \equiv \hat{A}$ is isometric when restricted to vectors of the form $\exp(i\Phi(f_0))$, since the time-reflection θ acts trivially on these vectors. It remains to see that the restriction to such vectors is *onto* \mathcal{H} . Then we wish to prove

$$(\mathcal{E}_0)^\wedge \supset \left(\bigcup_{t>0} \mathcal{E}_t \right)^\wedge. \quad (4.14)$$

First, let us see why (4.14), if true, finishes the proof. We must show that $\bigcup_{t>0} \mathcal{E}_t$ is dense in \mathcal{E}_+ . Of course, \mathcal{E}_+ is spanned by polynomials in classical fields of the form

$$\Phi(f) = \int \Phi(x, t) f(x, t) \sqrt{\mathcal{G}} dx dt.$$

Write the t integral as a Riemann sum:

$$\Phi(f) = \lim_{N \rightarrow \infty} \sum_{i=1}^N (\delta t)_i \Phi((f_i)_{t_i}), \quad (4.15)$$

$$\text{where } \Phi((f_i)_{t_i}) = \int \Phi(x, t_i) f_i(x) \sqrt{\mathcal{G}} dx, \quad (4.16)$$

and where $f_i(x) = f(x, t_i)$.

Eqn. (4.15) represents $\Phi(f)$ as a limit of linear combinations of elements $\Phi(f_{t_i}) \in \mathcal{E}_{t_i}$. A similar argument applies to polynomials $A(\Phi)$ of classical fields, and to L^2 limits of such polynomials. Thus $\bigcup_{t>0} \mathcal{E}_t$ is dense in \mathcal{E}_+ . Then (4.14) implies $(\mathcal{E}_0)^\wedge$ is also dense in \mathcal{E}_+ .

Equation (4.14) is proved by means of the following identity:

$$\langle \hat{A}, : \exp(i\Phi(f_t)) : \wedge \rangle_{\mathcal{H}} = \langle \hat{A}, : \exp(i\Phi(f^t_0)) : \wedge \rangle_{\mathcal{H}} \quad (4.17)$$

where

$$f^t := (F^{-1/2} e^{-t\omega} F^{1/2}) f, \quad (4.18)$$

where f is a function on Σ , and hence so is f^t . Thus

$$f^t_0(p, t') = \delta(t') (F^{-1/2} e^{-t\omega} F^{1/2} f)(p) \quad \text{for } p \in \Sigma.$$

To prove (4.17), we first suppose $A = : e^{i\Phi(g_s)} :$ where $g \in \mathcal{T}_\Sigma$ and $s > 0$. Then

$$\begin{aligned} \langle \hat{A}, : \exp(i\Phi(f_t)) : \rangle_{\mathcal{H}} &= \langle : e^{i\Phi(\theta g_s)} :, : e^{i\Phi(f_t)} : \rangle_{\mathcal{E}} \\ &= \exp \langle \theta g_s, C f_t \rangle_M \\ &= \exp \langle g, F^{1/2} \frac{e^{-(t+s)\omega}}{2\omega} F^{1/2} f \rangle_{\mathcal{K}_\Sigma} \end{aligned} \quad (4.19)$$

where we have used localization (Theorem 4.3) in the last line.

Computing the right side of (4.17) gives

$$\begin{aligned} \langle : e^{i\Phi(\theta g_s)} :, : e^{i\Phi(f^t_0)} : \rangle_{\mathcal{E}} &= \exp \langle \theta g_s, C(f^t_0) \rangle_M \\ &= \exp \langle g, F^{1/2} \frac{e^{-s\omega}}{2\omega} F^{1/2} f^t \rangle_{\mathcal{K}_\Sigma} \\ &= \exp \langle g, F^{1/2} \frac{e^{-(t+s)\omega}}{2\omega} F^{1/2} f \rangle_{\mathcal{K}_\Sigma} = (4.19). \end{aligned}$$

We conclude that eqns. (4.17)-(4.18) hold true for $A = : e^{i\Phi(g_s)} :$. We then infer the validity of (4.17) for all A in the span of $\bigcup_{t>0} \mathcal{E}_t$ by linear combinations and limits.

Equation (4.17) says that for every vector v in a set that is dense in \mathcal{H} , there exists $v' \in (\mathcal{E}_0)^\wedge$ such that $L(v) = L(v')$ for any linear functional L on \mathcal{H} . If $v \neq v'$ then we could find some linear functional to separate them, so they are equal. Therefore $(\mathcal{E}_0)^\wedge$ is a dense set, completing the proof of Theorem 4.4.

Theorem 4.4 implies that the physical Hilbert space is isometrically isomorphic to \mathcal{E}_0 , and to an L^2 space of the Gaussian measure with covariance which can be found by the $t, s \rightarrow 0$ limit of (4.19), to be:

$$\mathcal{H} = L^2(N_{d-1}^*, d\phi_{\mathfrak{C}}), \quad \text{where } \mathfrak{C} = F^{1/2} \frac{1}{2\omega} F^{1/2}, \quad (4.20)$$

and N_{d-1} denotes the nuclear space over the $(d-1)$ -dimensional slice. Compare (4.20) with [23], eqn. (6.3.1). By assumption, 0 lies in the resolvent set of ω , implying that \mathfrak{C} is a bounded, self-adjoint operator on \mathcal{K}_Σ .

4.5 The φ bound

Here we prove that an estimate known in constructive field theory as the *Glimm-Jaffe φ bound* (see [20]) is also true for curved spacetimes.

Theorem 4.5 (φ bound) *Let $T > 0$. There exists a constant M such that*

$$\left\langle \hat{A}, e^{-(H_0 + \varphi(h))T} \hat{A} \right\rangle_{\mathcal{H}} \leq \exp\left(T \|h\|_G^2 M\right) \|\hat{A}\|_{\mathcal{H}}^2, \quad (4.21)$$

where $\|h\|_G = \langle h, Gh \rangle^{1/2}$ and G is the resolvent of Q at $-m^2$.

Proof Apply the Schwartz inequality (for the inner product on \mathcal{H}) n times, to obtain

$$\begin{aligned} \left\langle \hat{A}, e^{-(H_0 + \varphi(h))T} \hat{A} \right\rangle_{\mathcal{H}} &\leq \|\hat{A}\|_{\mathcal{H}} \left\langle \hat{A}, e^{-2T(H_0 + \varphi(h))} \hat{A} \right\rangle_{\mathcal{H}}^{1/2} \\ &\leq \|\hat{A}\|_{\mathcal{H}}^{2-2^{-(n-1)}} \left\langle \hat{A}, e^{-2^n T(H_0 + \varphi(h))} \hat{A} \right\rangle_{\mathcal{H}}^{2^{-n}}. \end{aligned}$$

Apply the Feynman-Kac formula to the very last expression, to obtain

$$\left\langle \hat{A}, e^{-(H_0 + \varphi(h))T} \hat{A} \right\rangle_{\mathcal{H}} \leq \|\hat{A}\|_{\mathcal{H}}^{2-2^{-(n-1)}} \left\langle \Theta A, e^{-\int_0^{2^n T} \Phi(h,t) dt} U(2^n T) A \right\rangle_{\mathcal{E}}^{2^{-n}}.$$

We cannot take the $n \rightarrow \infty$ limit at this point, because the object depends on A . It suffices to establish the desired result for A in a dense subspace, so take $A \in L^4 \cap \mathcal{E}_+$. We now use the Schwartz inequality on \mathcal{E} as well as the fact that Θ is unitary on \mathcal{E} , to obtain

$$\left\langle \hat{A}, e^{-(H_0 + \varphi(h))T} \hat{A} \right\rangle_{\mathcal{H}} \leq \|\hat{A}\|_{\mathcal{H}}^{2-2^{-(n-1)}} \|A\|_{\mathcal{E}}^{2^{-n}} \left\langle A, e^{2 \int_0^{2^n T} \Phi(h,t) dt} A \right\rangle_{\mathcal{E}}^{2^{-(n+1)}}$$

Now Hölder's inequality with exponents $\frac{1}{4} + \frac{1}{4} + \frac{1}{2} = 1$ implies

$$\begin{aligned} \left\langle \hat{A}, e^{-(H_0 + \varphi(h))T} \hat{A} \right\rangle_{\mathcal{H}} &\leq \|\hat{A}\|_{\mathcal{H}}^{2-2^{-(n-1)}} \|A\|_{\mathcal{E}}^{2^{-n}} \|A\|_{L^4}^{2^{-n}} \left(\int e^{4 \int_0^{2^n T} \Phi(h,t) dt} d\mu_0 \right)^{2^{-n-2}}. \quad (4.22) \end{aligned}$$

Up to this point, the argument applies to a general measure $d\mu$ on path space. Now assume that the measure is Gaussian. The function $f = 4h(\mathbf{x})\chi_{[0, 2^n T]}(t)$ has the desirable property that $\Phi(f) = 4 \int_0^{2^n T} \Phi(h, t) dt$, so the Gaussian integral in (4.22) equals $S(if) = e^{\langle f, Cf \rangle / 2}$. Therefore,

$$\left\langle \hat{A}, e^{-(H_0 + \varphi(h))T} \hat{A} \right\rangle_{\mathcal{H}} \leq \|\hat{A}\|_{\mathcal{H}}^{2-2^{-(n-1)}} \|A\|_{\mathcal{E}}^{2^{-n}} \|A\|_{L^4}^{2^{-n}} S(if)^{2^{-n-2}}. \quad (4.23)$$

For H_1 and H_2 self-adjoint operators with $0 \leq H_1 \leq H_2$, we have $(H_2 + a)^{-1} \leq (H_1 + a)^{-1}$ for any $a > 0$. By theorem 4.2, $-Q \geq 0$, so take $H_1 = -Q$, and $H_2 = -(1/F)\partial_t^2 - Q$. We conclude³

$$C = (-\Delta + m^2)^{-1} \leq (-Q + m^2)^{-1} \equiv G.$$

³ Compare this with the analogous estimate valid in \mathbb{R}^d , $C \leq (-\nabla_{\mathbf{x}}^2 + m^2)^{-1}$, which may be proved by a Fourier transform of the resolvent kernel.

Since $\ker(G) = \{0\}$, G determines a norm $\|h\|_G = \langle h, Gh \rangle^{1/2}$. Then

$$S(if) \leq e^{8\langle h, Gh \rangle^{2nT}} = e^{2^{n+3}T\|h\|_G^2}.$$

Raising this to the power 2^{-n-2} , and taking the $n \rightarrow \infty$ limit we see that the factors $\|A\|_{\mathcal{E}}^{2^{-n}} \|A\|_{L^4}^{2^{-n}}$ approach 1, and thus (4.23) becomes:

$$\left\langle \hat{A}, e^{-(H_0 + \varphi(h))T} \hat{A} \right\rangle_{\mathcal{H}} \leq e^{2T\|h\|_G^2} \|\hat{A}\|_{\mathcal{H}}^2.$$

This establishes (4.21), completing the proof of Theorem 4.5.

4.6 Fock representation for time-zero fields

To obtain a Fock representation of the *time-zero fields* we mimic the construction of [23, § 6.3] with the covariance (4.20).

To simplify the constructions in this section, we assume the form $ds^2 = dt^2 + G_{\mu\nu} dx^\mu dx^\nu$ and $F = 1$. Then $Q = \Delta_\Sigma$, the Laplacian on the time-zero slice, and $\mu = (-\Delta_\Sigma + m^2)^{1/2}$. The set of functions $h \in L^2(\Sigma)$ such that $\mu^p h \in L^2(\Sigma)$ is precisely the Sobolev space $H_p(\Sigma)$, which is also the set of h such that $\mathfrak{C}^{-p} h \in L^2$. Sobolev spaces satisfy the reverse inclusion relation $p \geq q \Rightarrow H_q \subseteq H_p$. Also $\mathfrak{C}^q f \in H_p \Leftrightarrow f \in H_{q-p}$.

This allows us to determine the natural space of test functions for the definition of the Fock representations:

$$\begin{aligned} a(f) &= \frac{1}{2} \phi \left(\mathfrak{C}^{-1/2} f \right) + i\pi \left(\mathfrak{C}^{1/2} f \right) \\ a^*(f) &= \frac{1}{2} \phi \left(\mathfrak{C}^{-1/2} f \right) - i\pi \left(\mathfrak{C}^{1/2} f \right). \end{aligned}$$

In particular, if the natural domain of ϕ is H_{-1} as discussed following eqn. (1.3), then f must lie in the space where $\mathfrak{C}^{-1/2} f \in H_{-1}$, i.e. $f \in H_{1/2}$.

5 Conclusions and Outlook

We have successfully generalized Osterwalder-Schrader quantization and several basic results of constructive field theory to the setting of static spacetimes.

Dimock [14] constructed an interacting $\mathcal{P}(\varphi)_2$ model with variable coefficients, with interaction density $\rho(t, x) : \varphi(x)^4$, and points out that a Riemannian $(\varphi^4)_2$ theory may be reduced to a Euclidean $(\varphi^4)_2$ theory with variable coefficients. However, the main constructions of [14] apply to the Lorentzian case and for curved spacetimes no analytic continuation between them is known. Establishing the analytic continuation is clearly a priority. Also, there are certain advantages to a perspective which remembers the spacetime structure; for example, in this picture the procedure for quantizing spacetime symmetries is more apparent.

In the present paper we have not treated the case of a non-linear field, though all of the groundwork is in place. Such construction would necessarily involve a generalization of the Feynman-Kac integral (2.11) to curved space, and would have far-reaching implications, and one would like to establish properties of the particle spectrum for such a theory.

The treatment of symmetry in this paper is only preliminary. We have isolated two classes of isometries, the reflected and reflection-invariant isometries, which have well-defined quantizations. We believe that this construction can be extended to yield a unitary representation of the isometry group, and work on this is in progress. This, together with suitable extensions of section 2.6 could have implications for the representation theory of Lie groups, as is already the case for the geometric quantization of classical Hamiltonian systems.

The treatment of variation of the metric in section 3 is also preliminary; it does not cover the full class of static spacetimes. Geroch [19] gave a rigorous definition of the limit of a family of spacetimes, which formalizes the sense in which the Reissner-Nordström black hole becomes the Schwarzschild black hole in the limit of vanishing charge. It would be interesting to combine the present framework with Geroch's work to study rigorously the properties of the quantum theory under a limit of spacetimes.

Another direction is to isolate specific spacetimes suggested by physics which have high symmetry or other special properties, and then to extend the methods of constructive field theory to obtain mathematically rigorous proofs of such properties. Several studies along these lines exist [7, 26], but there is much more to be done. We hope that the Euclidean functional integral methods developed here may facilitate further progress. Rigorous analysis of thermal properties such as Hawking radiation should be possible. Given that new mathematical methods are available which pertain to Euclidean quantum field theory in AdS, a complete, rigorous understanding of the holographically dual theory on the boundary of *AdS* suggested by Maldacena [1, 24, 36, 45] may be within reach of present methods.

Constructive field theory on flat spacetimes has been developed over four decades and comprises thousands of published journal articles. Every statement in each of those articles is either: (i) an artifact of the zero curvature and high symmetry of \mathbb{R}^d or \mathbb{T}^d or (ii) generalizable to curved spaces with less symmetry. The present paper shows that the Osterwalder-Schrader construction and many of its consequences are in class (ii). For each construction in class (ii), investigation is likely to yield non-trivial connections between geometry, analysis, and physics.

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A Euclidean Anti-de Sitter and its Analytic Continuation

The Green's function G on a general curved manifold is the inverse of the corresponding positive transformation, so it satisfies

$$(\Delta - \mu^2)G = -g^{-1/2}\delta, \quad (\text{A.1})$$

where $G(p, q)$ is a function of two spacetime points. By convention Δ acts on G in the first variable, and δ denotes the Dirac distribution of the geodesic distance $d = d(p, q)$. Translation invariance implies that G only depends on p and q through $d(p, q)$. We note that solutions of the homogeneous equation $(\Delta - \mu^2)\phi = 0$ may be recovered from the Green's function. Conversely, we may deduce the Green's function by solving the homogeneous equation for $d > 0$ and enforcing the singularity at $d = 0$.

The equation (A.1) for the Green's function takes a simple form in geodesic polar coordinates on \mathbb{H}^n with $r = d =$ geodesic distance; the Green's function has no dependence on the angular variables and the radial equation yields

$$(\partial_r^2 + (n-1)\coth(r)\partial_r - \mu^2)G(r) = -\delta(r). \quad (\text{A.2})$$

We find it convenient to write the homogeneous equation in terms of the coordinate $u = \cosh(r)$. When $u \neq 1$, (A.2) becomes

$$(\Delta - \mu^2)G(u) = -(1-u^2)G''(u) + nuG'(u) - \mu^2G(u) = 0. \quad (\text{A.3})$$

For $n = 2$ and $\mu^2 = \nu(\nu+1)$, eqn. (A.3) is equivalent to Legendre's differential equation:

$$(1-u^2)Q''_\nu(u) - 2uQ'_\nu(u) + \nu(\nu+1)Q_\nu(u) = 0. \quad (\text{A.4})$$

Note that (A.4) has two independent solutions for each ν , called Legendre's P and Q functions, but the Q function is selected because it has the correct singularity at $r = 0$. Thus

$$G_2(r; \mu^2) = \frac{1}{2\pi}Q_\nu(\cosh r), \quad \text{where } \nu = -\frac{1}{2} + \left(\mu^2 + \frac{1}{4}\right)^{1/2}. \quad (\text{A.5})$$

The case $\mu^2 = 0$ is particularly simple; there the Legendre function becomes elementary:

$$G_2(r; 0) = -\frac{1}{2\pi} \ln\left(\tanh \frac{r}{2}\right) = \frac{1}{2\pi}Q_0(\cosh r). \quad (\text{A.6})$$

For $n = 3$, one has

$$G_3(r; \mu^2) = \frac{1}{4\pi} \frac{e^{\pm r\sqrt{\mu^2+1}}}{\sinh(r)}. \quad (\text{A.7})$$

Finally, we note that the analytic continuation of (A.5) gives the Wightman function on AdS_2 . The real-time theory on Anti-de Sitter, including its

Wightman functions, were discussed by Bros et al. [7]. In particular, our equation (A.5) analytically continues to their equation (6.8).

Given a complete set of modes, one may also calculate the Feynman propagator by using the relation $iG_F(x, x') = \langle 0 | T\{\phi(x)\phi(x')\} | 0 \rangle$ and performing the mode sum explicitly as in [9]; the answer may be seen to be related to the above by analytic continuation. Here, T denotes an AdS -invariant time-ordering operator. A good general reference is the classic paper [3].

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